

External Exposure to Natural Radiation

Results of some Experimental Studies

1- INRNE/NPI collaboration

Content

- Comparison of methods used to characterise the exposure due to the environmental radiation background
- Studies on the territory of the INRNE
- Studies on the high-mountain station at Moussala, comparison with Lomnický Štít (High Tatras)
- Common studies with TLDs (aircraft, HECF)

Methods used

1. Active instruments:

- ***Environmental radiation dose rate meter NB 3201*** with a plastic scintillator as sensitive element; able to measure the environmental radiation background with low linear energy transfer (LET) (10 nSv to few mSv/h)
- ***MDU-Liulin semiconductor spectrometer*** with Si-diode as the sensitive element; able to estimate both low LET and neutron component of the natural background (10 nSv to few mSv/h)

2. Passive detectors:

- ***Thermoluminescent detectors (TLD) $\text{CaSO}_4\text{:Dy}$***
- ***Moderator sphere (12 inches) with TED in contact with B-radiators***

Measuring localities

1. *Territory of the INRNE BAS at Sofia:*

- Outside of the building of the Division of Cosmic Ray Physics (DCRP) and on its terrace
- Around the reactor IRT 2000 building

2. *Basic Ecological Observatory (BEO)*

Moussala; altitude 2925 m, geographic coordinates 25°35' E and 42°11' N

3. *Lomnický Štít Observatory of the Institute of Experimental Physics SAV, High Tatras*; altitude 2634 m, geographic coordinates 20°22' E and 49°20' N

Results - INRNE

Measuring point	H*(10), nSv/h, as measured by		
	NB 3201	MDU-Liulin*	TLD
DCRP – in front of	106	-	-
DCRP – terrace	102	103	78-141
Reactor – 1	110	134	130-154
Reactor - 2	100	104	124

* Mean value for two MDU units

Results – BEO Moussala

Measuring point	H*(10), nSv/h, as measured by		
	NB 3201	MDU	TLD
Ground level of building	240	-	-
First floor of building	138	130 ^{*)}	-
Outside of building	176	-	165-183
Terrace of building	138	140 ^{*)}	-

***) Corresponding only to low LET radiation**

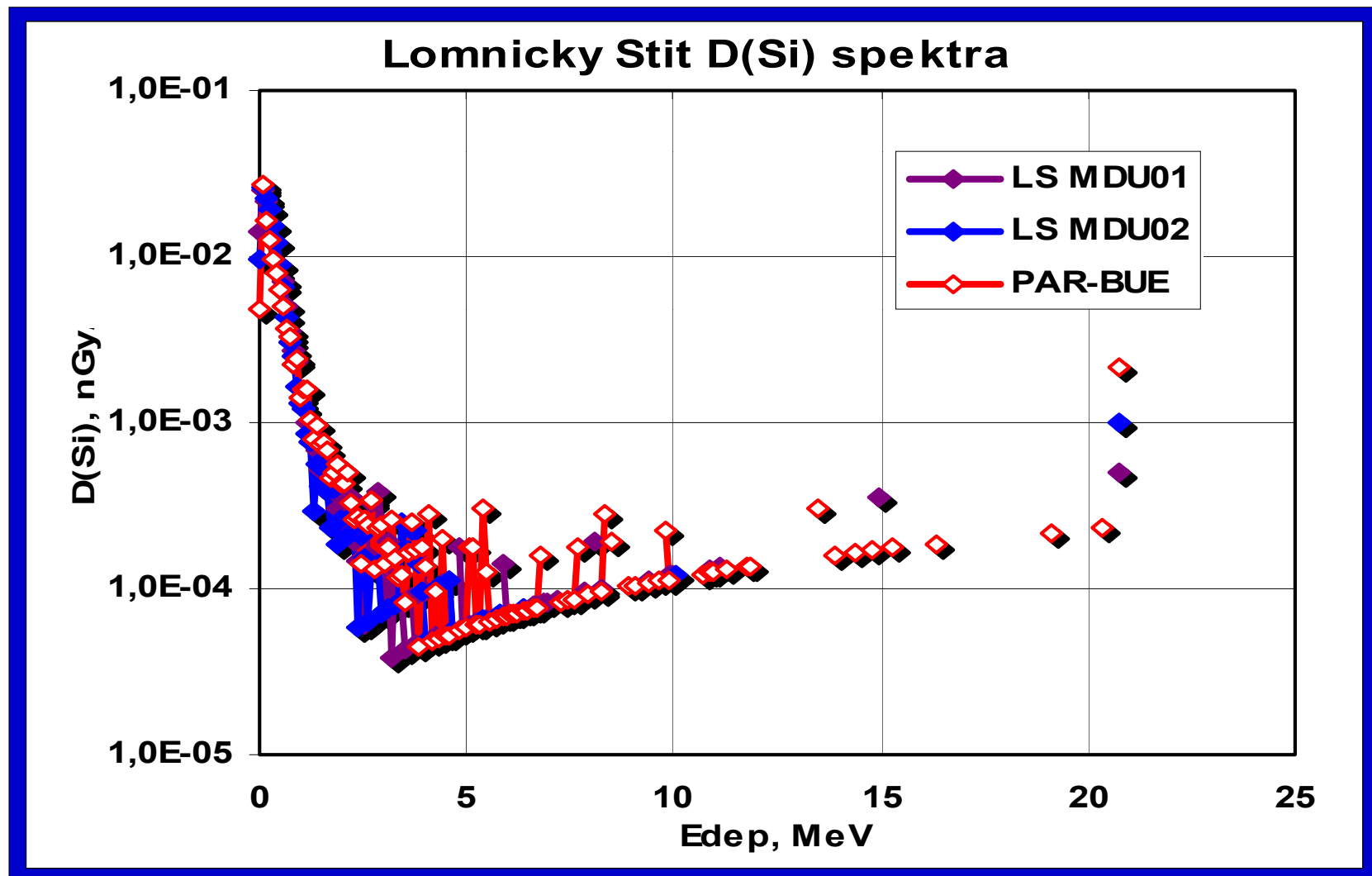
**The average value of monitors of INRNE BAS110 nSv/h;
minimum and maximum values 87.5 and 128.1 mSv/h**

Results - neutron component of cosmic radiation at BEO Moussala - Comparison of different instruments and methods

Method (Quantity measured)	Annual value measured	Annual value corrected
Harwell 3208-1	$(190 - 390) \mu\text{Sv}^{1)}$	$(380 - 780) \mu\text{Sv}$
Sphere with ^{10}B	$(254 \pm 26) \mu\text{Sv}^{2)}$	$(508 \pm 52) \mu\text{Sv}$
MDU-Liulin	$(180 \pm 72) \text{nGy}^{3)}$	$(1080 \pm 380) \mu\text{Sv}$
UNSCEAR 2000	-	$\sim 580 \mu\text{Sv}$

¹⁾ Period 09/02 ÷ 05/03; ²⁾ Period 11/00 ÷ 02/04; ³⁾ November 2000

Results - neutron component of cosmic radiation at Lomnický Štít - Event spectra in MDU Si-diode instrument



Results - neutron component of cosmic radiation- Comparison of BEO Moussala and Lomnický Štít high mountain stations

Estimated from	Annual values measured at			
	Moussala ¹⁾		Lom. Štít ¹⁾	
	D(Si), μGy	H*(10), μSv	D(Si), μGy	H*(10), μSv
MDU-Fr	180 ± 72	1080 ± 380	-	-
MDU-CZ01	-	-	135 ± 6	440 - 850
MDU-CZ02	-	-	113 ± 7	440 - 760
UNSCEAR 2000	-	~ 580	-	~ 570

¹⁾ Measuring times 14 hours at Moussala, 95 hours at Lomnický Štít

Onboard aircraft exposure

- A310-300 Czech Airlines
- 05/05-31/12/05
- Together with MDU-Liulin
- All navigation data available – dose calculated
- Trach etched detectors; TLD's

TLD results:

1. Exposure dates

- NPI's – full time
- INRNE – since 08/09

2. Results since 08/09/05:

$\text{Al}_2\text{O}_3 : \text{C}$: (1.75 ± 0.25) mSv

AIP glas: (1.66 ± 0.09) mSv

$\text{CaSO}_4 : \text{Dy}$: $(1.5-1.8)$ mSv

RR of TLDs to high energy charged particles

HIMAC, NIRS Chiba, September 2005

Ion	Energy (MeV/n)	LET keV/ μ m	shield	RR		
				Al ₂ O ₃ /C	Al-P skla	CaSO ₄
⁴ He	150	2,2	Bare Beam	0,77	0,84	1,01
⁴ He	150	3,5	Behind PMMA	1,10	1,37	1,63
¹⁶ O	400	19,8	Bare Beam	0,50	0,98	0,99
¹⁶ O	400	23,8	Behind PMMA	0,47	0,95	0,94
⁴⁰ Ar	500	89,3	Bare Beam	0,34	0,68	0,62
⁴⁰ Ar	500	100	Behind PMMA	0,34	0,68	0,62
⁵⁶ Fe	200	310	Bare Beam	0,26	0,61	0,47
⁵⁶ Fe	200	321	Behind PMMA	0,35	0,77	0,57

External Exposure to Natural Radiation

Results of some Experimental Studies

2 - Onboard aircraft exposure

1. International activities

2. Experimental studies 1991-2005

- **Period 1991-1999 – set of dosimetry methods (TEPC, rem-meter, Bubbles, TLD, scintillator, IC, GM, APD)**
- **Period since 2000 – mostly MDU-Liulin results**

3. Routine individual dosimetry of aircraft crew members of czech air companies, 1998-2004

ICRU&ICRP TG No. 50

- **Goal : to prepare the recommendations
“Reference doses from cosmic radiation
exposure of aircraft crew”**
- **Scheduled: 2001-2005; now 2006/2007**
- **Current stage: decision taken to concile
different groups and prepare the data in
the forme of acceptable limits for low and
high „typical“ solar activities periods**

WG 21 ISO TC/85 - since 2001

- **ISO 20785: “Dosimetry for exposure to cosmic radiation in civilian aircraft”**
- **Current stage: decided to prepare in 3 parts:**
 - **“Conceptual basis for measurements”;**
done - 09/05
 - **“Characterisation of instrument response”,**
WD – end 2005
 - **“Standardization of measurements procedures” (≥ 2006)**

DG's RTD and TREN EC + EURADOS

- **WG – Article 31 EURATOM**
- **DOSMAX (6thEF)**
- **CONRAD – WP 6 (7thEF)**
- **EURADOS WG 5**

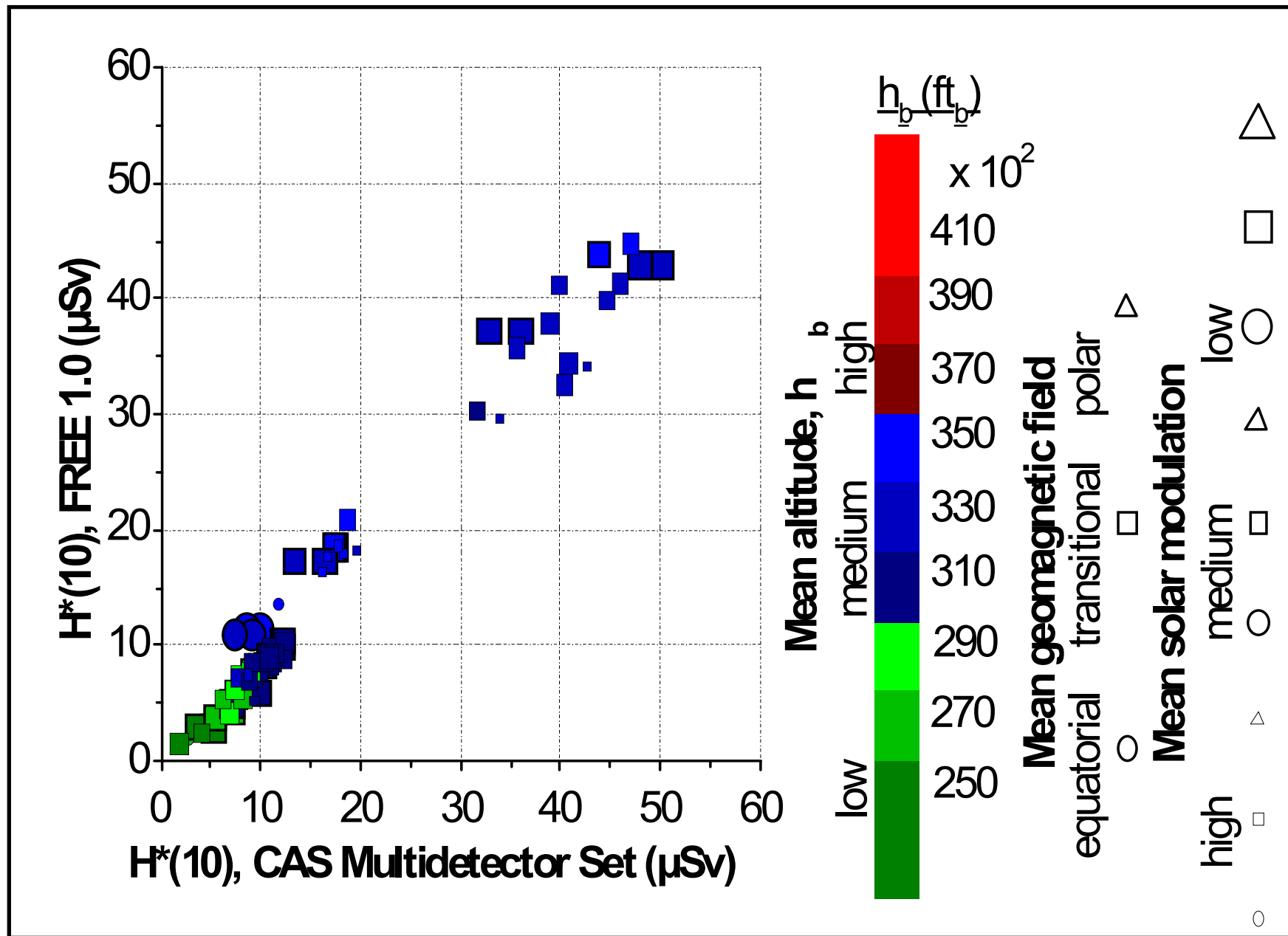
WG EURADOS (Art. 31 EURATOM)

„Cosmic Radiation Exposure of Aircraft Crew: Compilation of available experimental and calculated data“

1999-2004; Eds.: I. McAulay, D. Bartlett, P. Beck,
K. Schnuer, H. Schraube, F. Spurny; 280 p.

ISBN 92-894-8448-9; several ten thousands

data: *The total uncertainty of measured
values is about 25% (2s). The total
uncertainty in the values of effective dose
calculated by different codes is estimated to
be about 30%.*



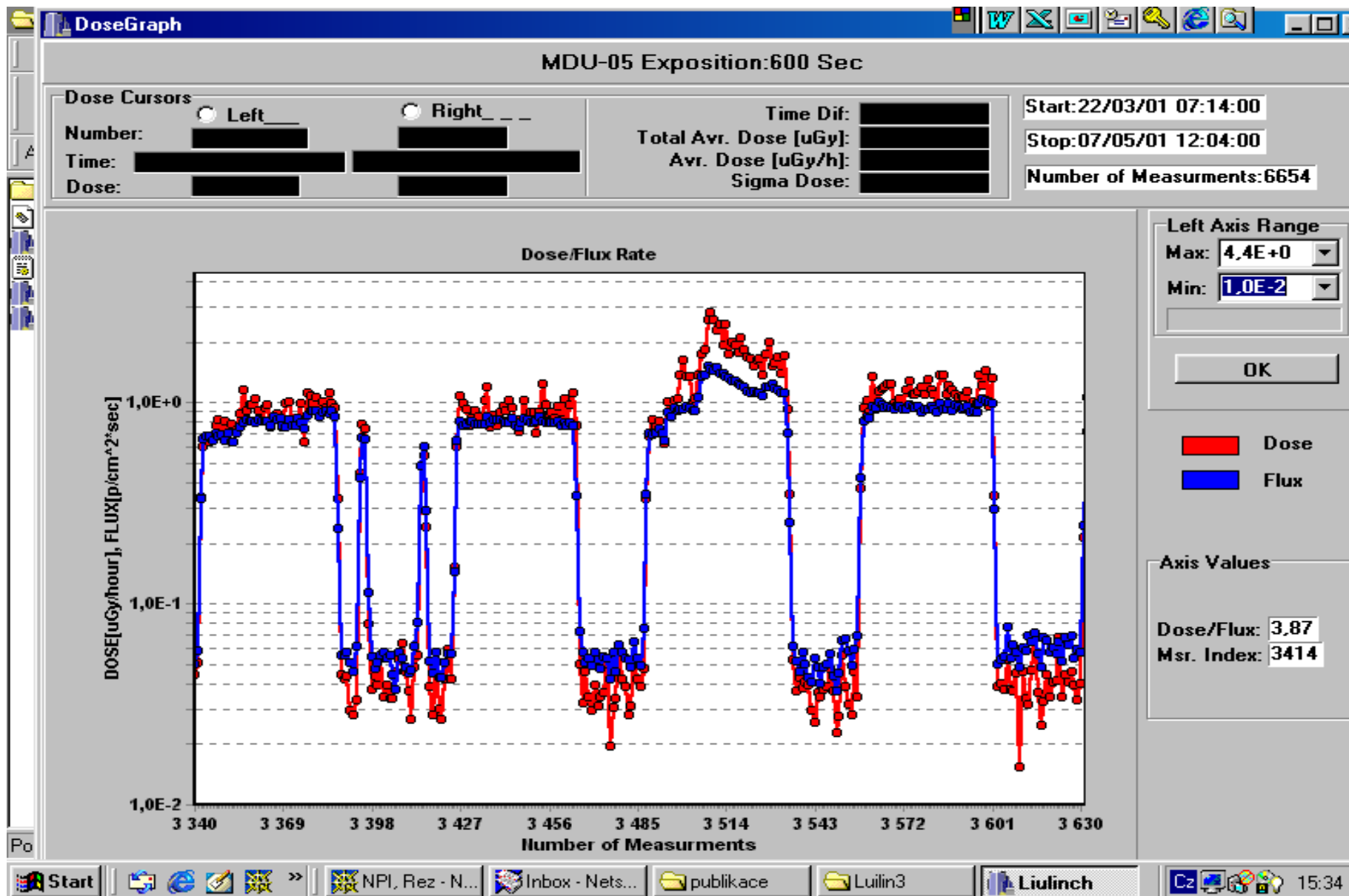
Example of results -long term monitoring-1

Period	Returned flights monitored^{*)}	Numbers of flights
22/03-07/05/01	PRG-NY(25), PRG-TOR(13), PRG-DUB(3)	108
30/05-24/07/01	PRG-NY(41), PRG-TOR(12)	125
29/08-16/10/01	PRG-NY(26), PRG-TOR(13), PRG-DUB(2)	96
25/10-10/12/01	PRG-NY(20), PRG-TOR(7), PRG-AMS(1)	70
06/05-28/06/02	PRG-NY(22), PRG-TOR(13), PRG-DUB(8), PRG-LHR(5), PRG-MAD(5)	124
16/10-06/12/02	PRG-NY(23), PRG-TOR(10), PRG-DUB(1), PRG-DUB-CM (5), PRG-MUN(1)	110

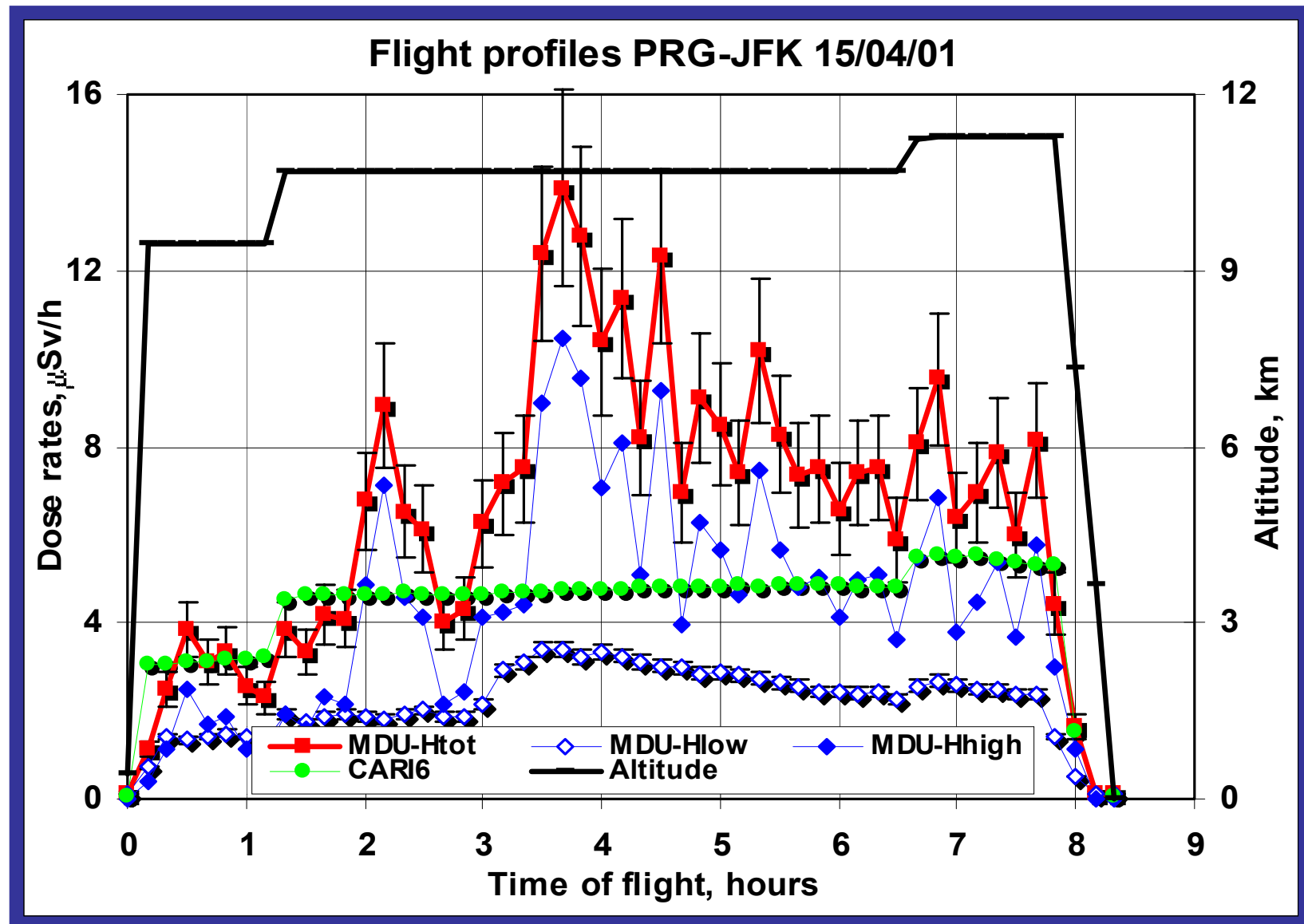
^{*)} PRG – Prague, NY – New York, TOR – Toronto (via Montreal), DUB – Dubai, CM-Colombo, MUN – Munchen, AMS – Amsterdam, LHR-London, MAD-Madrid

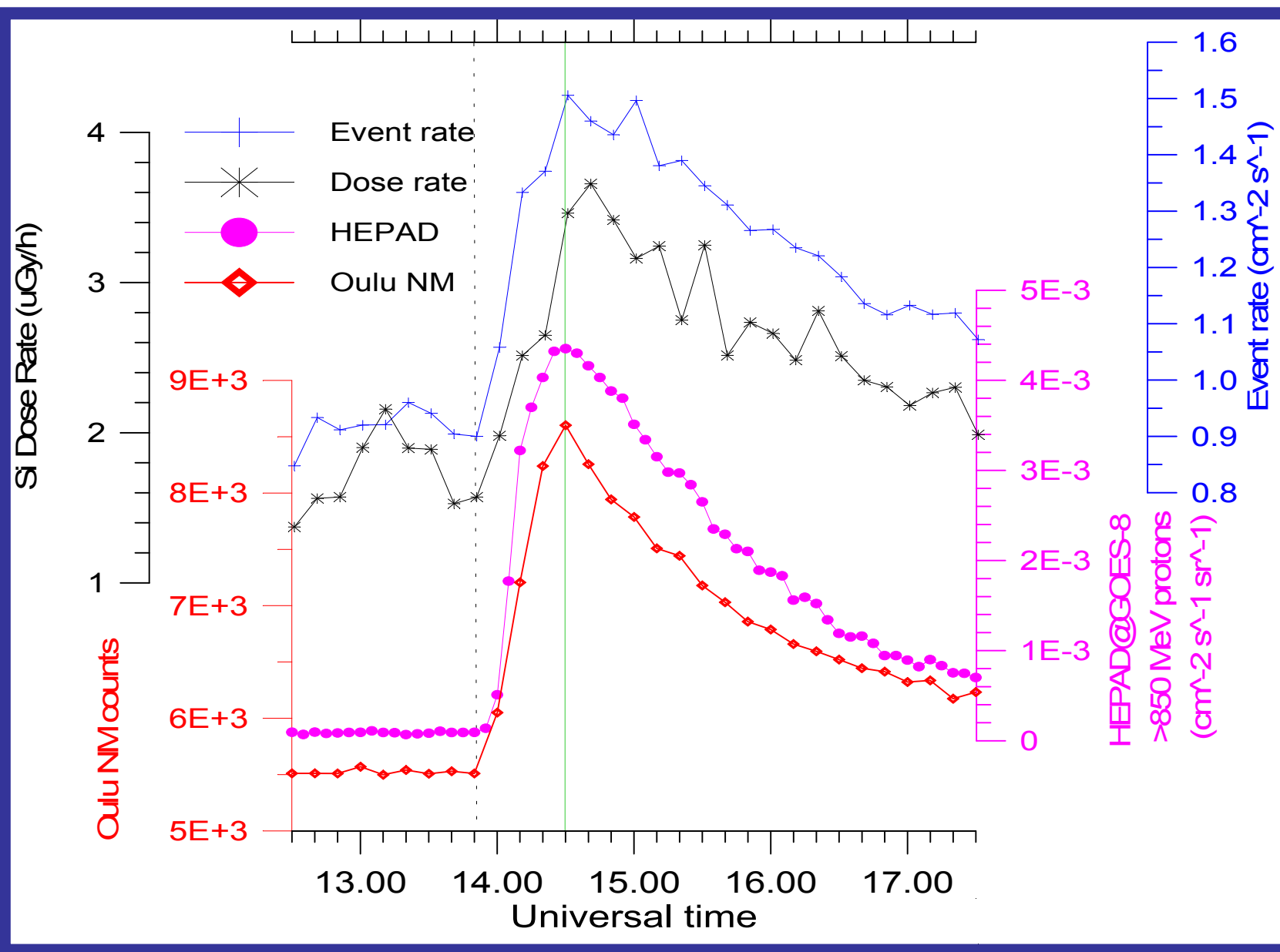
Example of results -long term monitoring-2

Route	Relative deviation, %, of measured and Epcard calculated H				
Period	22/03- 07/05/01	30/05- 24/07/01	29/08- 16/10/01	25/10- 10/12/01	16/10- 06/12/02
PRG - JFK	0.1±5.3	- 11.5±4.8	-5.7±3.7	-6.6±5.1	4.6±3.5
JFK - PRG	0.5±6.0	11.2±5.9	-5.6±4.0	-7.1±4.6	2.9±3.6
PRG - YYZ	-1.8	-9.1±1.5	-3.7±2.9	-	-
YYZ - PRG	9.8	-8.9±1.0	-5.2±3.6	-	-
PRG - YUL	2.6±6.9	-8.9±5.2	-5.2±3.6	-5.9±4.8	-0.4±5.3
YUL - PRG	0.6±3.8	-10.1±3.5	-0.5±5.0	-2.5±4.8	-0.8±4.8
PRG - DXB	-6.7±2.8	-	-13.7±0.1	-	9.2±3.3
DXB - PRG	-8.8±6.4	-	-6.8±2.7	-	13.4±4.2

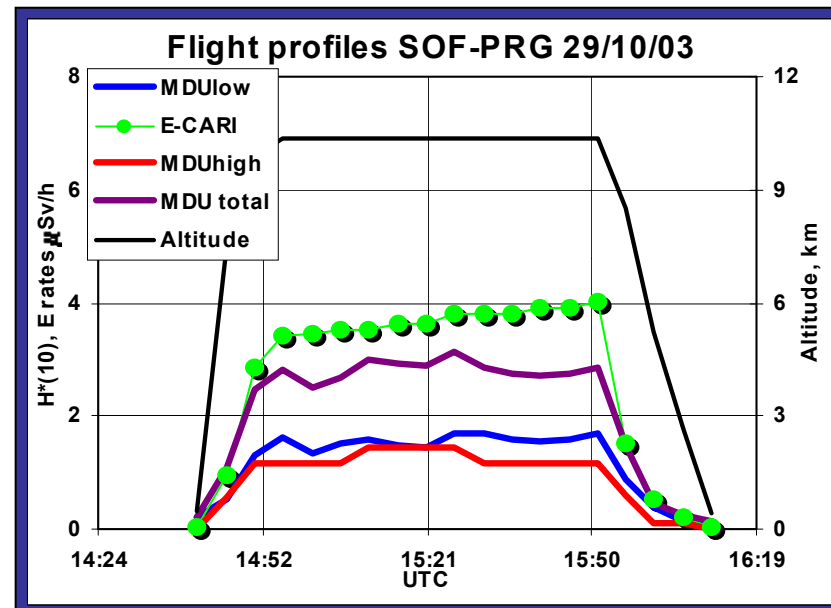
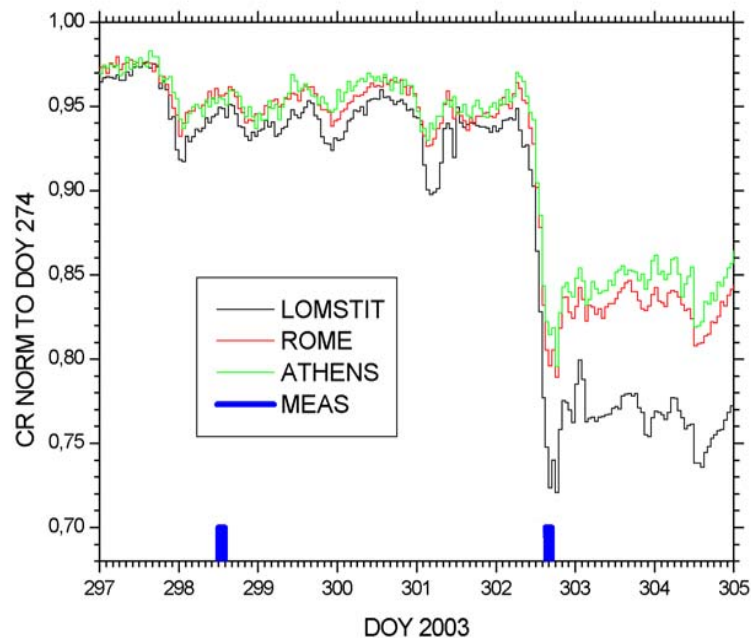
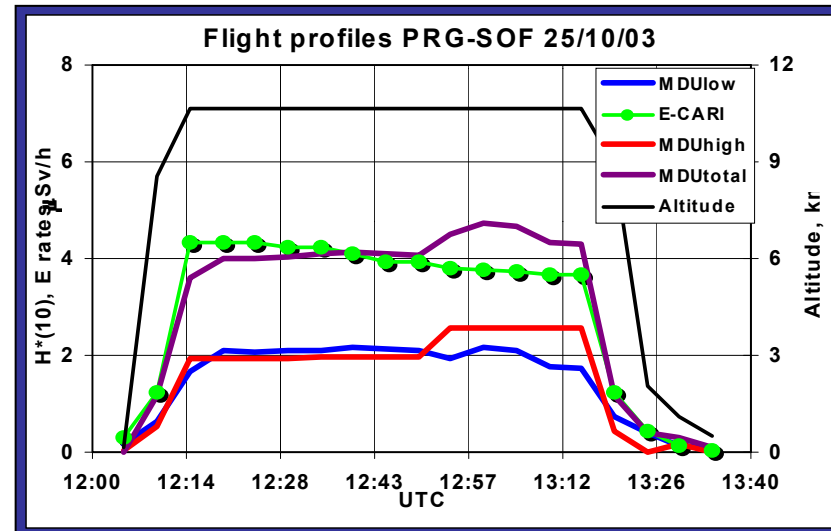
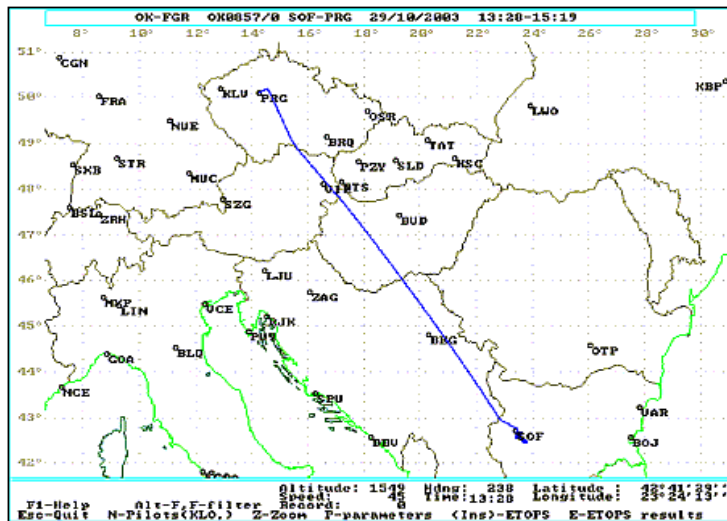


Solar flare (GLE 60) registered onboard with MDU aircraft; rate increased 3 times, total E by 40%





Registration of extreme events: Forbush decrease after GLE 65 29/10/2003



Ratios of ambient dose equivalent values for total flights estimated from MDU-Liulin measurements and calculated by means of EPCARD3.2 code

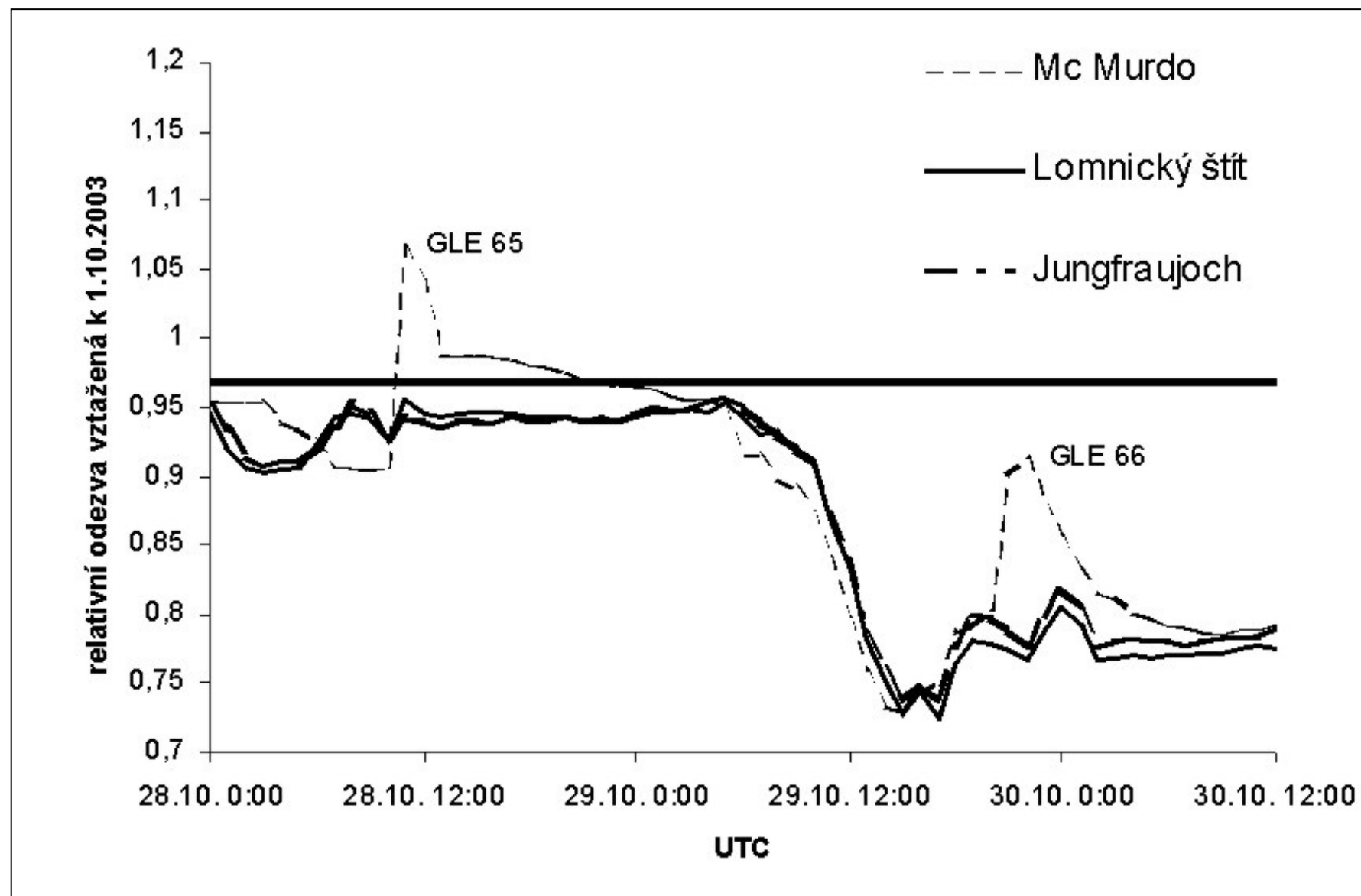
Date	Event	For low E_{dep}	For high E_{dep}	Total H
12/04/01	Forbush (LŠ: - 13 %)	0,92	0,78	0,84
15/04/01	GLE 60 (LŠ: + 10 %)	1,24	1,68	1,45
16/04/01	normal	1,01	1,00	1,01
25/10/03	normal	0,95	1,00	0,98
29/10/03	Forbush (LŠ: - 26 %)	0,88	0,62	0,74

Remarks:

- 1) LŠ: difference of countings at Lomnický Štít cosmic ray neutron monitor
- 2) April 2001: PRG-JFK flights; October 2003: PRG-SOF (25/10); SOF-PRG (29/10)

Cosmic ray neutron monitor stations

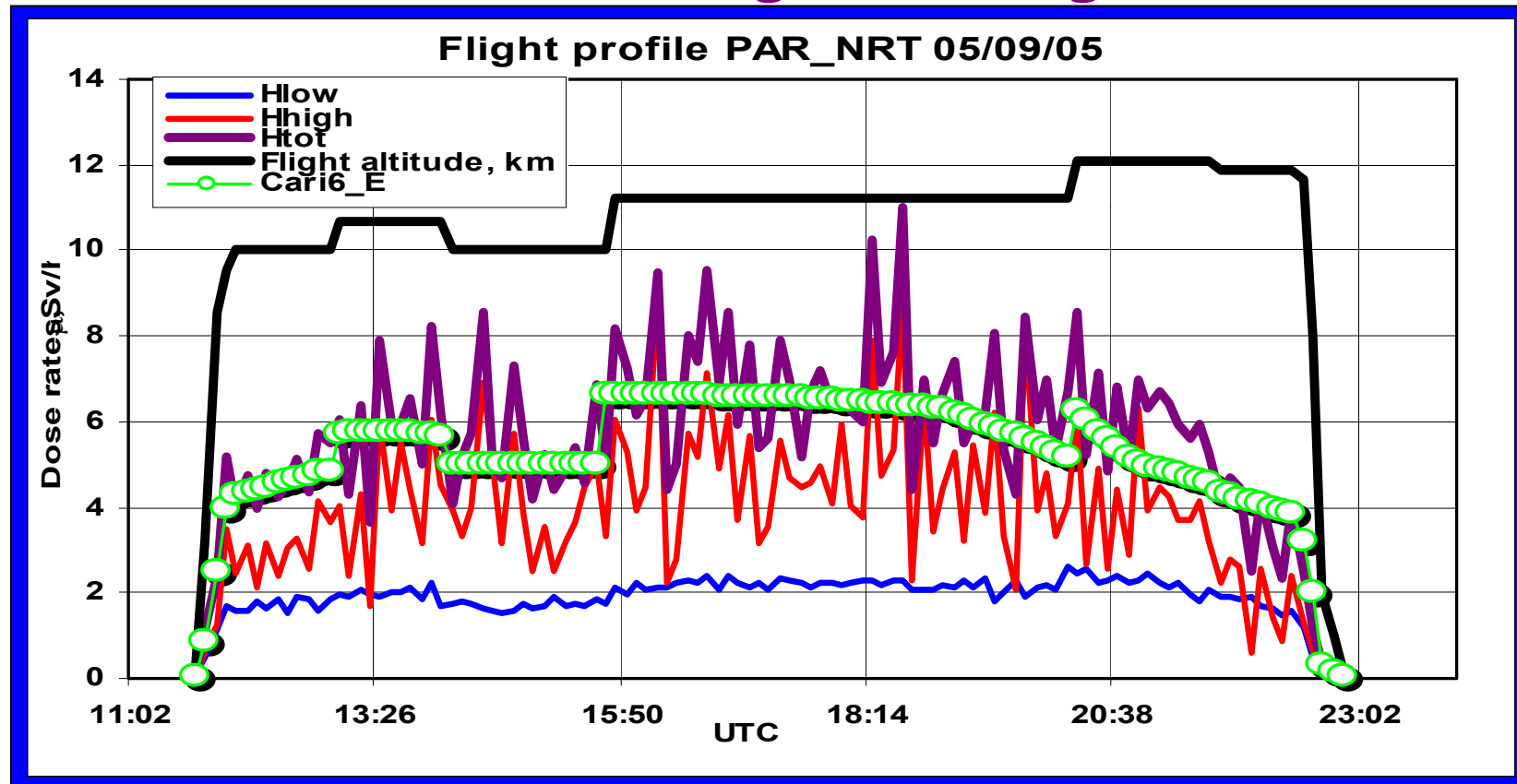
Monitor	Country	Latitude	Longitude	Altitude [m]	Rigidity [GV]
Jungfraujoch	Switzerland	46,55° N	7,98° E	3550	4,48
Lomnický štít	Slovakia	49,2° N	29,22° E	2632	3,84
McMurdo	Antarctic	77,85° S	166,72° E	48	0,01
Newark	USA	39,7° N	75,7° W	50	2,1
South Pole	Antarctic	90° S	0°	2820	0,09
Thule	Groenland	76,58° N	68,42° W	260	0



Measurements 2005

- Long-haul flights Prague Amsterdam-Beijing-Amsterdam-Prague 26/03 – 02/04;
- Long-haul flight Paris-Tokyo 05/09.
- Individual flights 2005;
- Long-term monitoring onboard a Czech Airlines aircraft

Other long-haul flight



Total exposure	MDU $H(E)_{app}$	CARI 6 E	EPCARDv.3.2 E	EPCARDv.3.2 $H^*(10)$
μSv	63.7 ± 8.6	59.0	56.1	47.0
relatively	1.08	1.00	0.95	0.80

Short-haul flights 2005

Flight	Date	MDU			CARI	EPC- E	EPC- H*(10)		
		Hlow	Hhigh	Htot			Htot	Hhigh	Hlow
PRG_SVO	17.II	3,52	5,51	9,03	8,73	8,5	7,5	4,2	3,3
SVO_PRG	22.II	2,84	5,95	8,79	7,38	7,5	6,5	3,7	2,8
PRG_SOF	9.III	2,12	2,21	4,33	4,41	4,5	4	2,16	1,84
SOF_PRG	15.III	1,97	3,01	4,98	4,67	4,5	4	2,16	1,84
PRG_AMS	26.III	1,28	2,36	3,64	3,56	3,48	2,48	1,42	1,06
AMS_PRG	2.IV	1,26	1,81	3,07	2,95	3,42	2,48	1,4	1,1
PRG_SOF	7.VII	1,76	2,67	4,43	4,37	5	4,5	2,43	2,17
SOF_PRG	13.VII	1,81	3,16	4,97	4,62	5	4,5	2,43	2,17
PRG_SVO	11.VIII	3,61	6,5	10,11	9,86	11,5	9,5	5,32	4,18
SVO_PRG	17.VIII	3,78	7,41	11,19	11,35	11,5	9,5	5,32	4,18
total		57,66	96,7	152,36	149,84	144,9	122,96	69,04	54,14

Average total value : $(142 \pm 13) \mu\text{Sv}$

Long-term monitoring onboard a CSA aircraft state-of-art 05/05/-31/12/05

- Total number of flights : ~ 400;
- More than 70 % - above North Atlantic;
- Interesting events:
 - Forbush decreases:
 - ✓ 8.-9.5.; (**PRG-JFK**)
 - ✓ 15.-16.5.; (**2 x PRG-YYZ; PRG-DXB-CMB**)
 - ✓ 17.-18.7. (small SPE ?) (**PRG-TUN (MAD)-PRG**)
 - ✓ 10.- 16.9. (**PRG-SVO; PRG-MAD; 2x PRG-JFK; AR**)

Onboard aircraft exposure

- A310-300 Czech Airlines
- 05/05-31/12/05
- Together with MDU-Liulin
- All navigation data available – dose calculated
- Trach etched detectors; TLD's

TLD results:

1. Exposure dates

- NPI's – full time
- INRNE – since 08/09

2. Results since 08/09/05:

$\text{Al}_2\text{O}_3 : \text{C} : (1.75 \pm 0.25) \text{ mSv}$

AIP glas: $(1.66 \pm 0.09) \text{ mSv}$

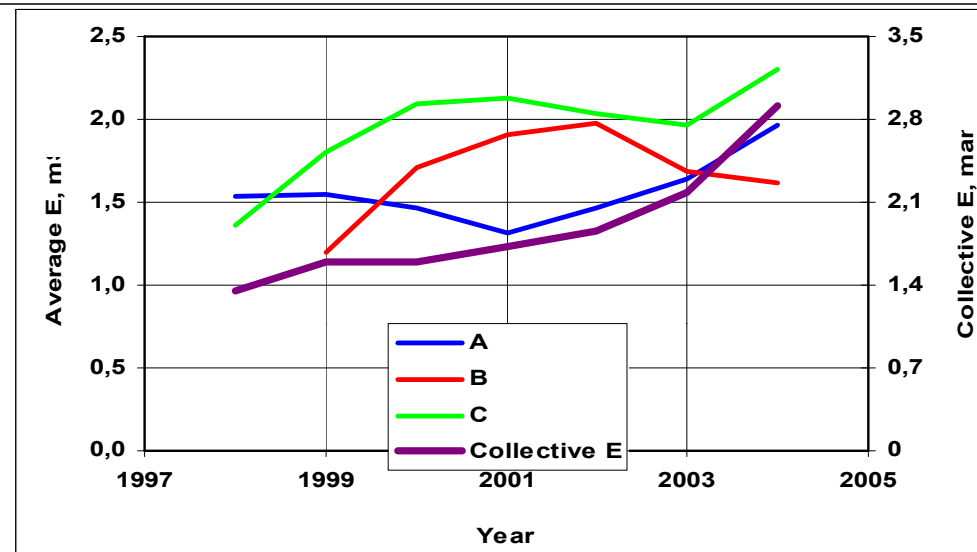
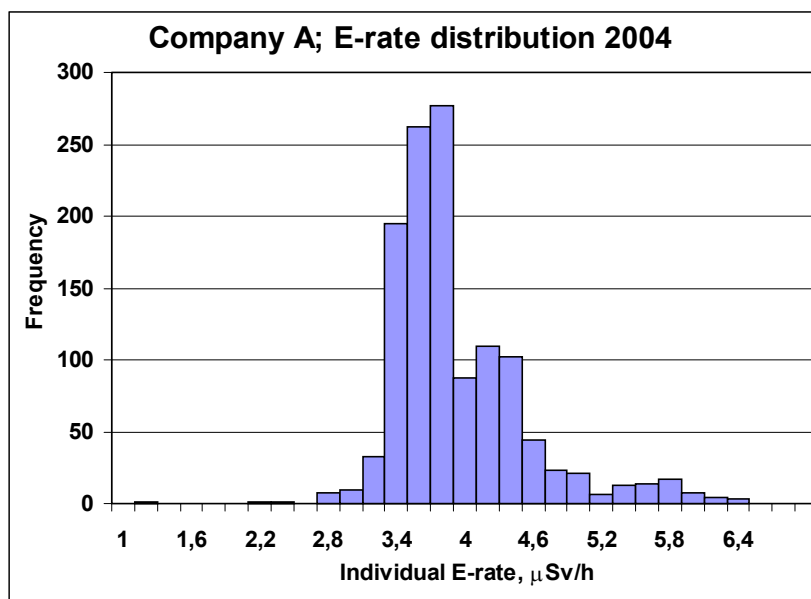
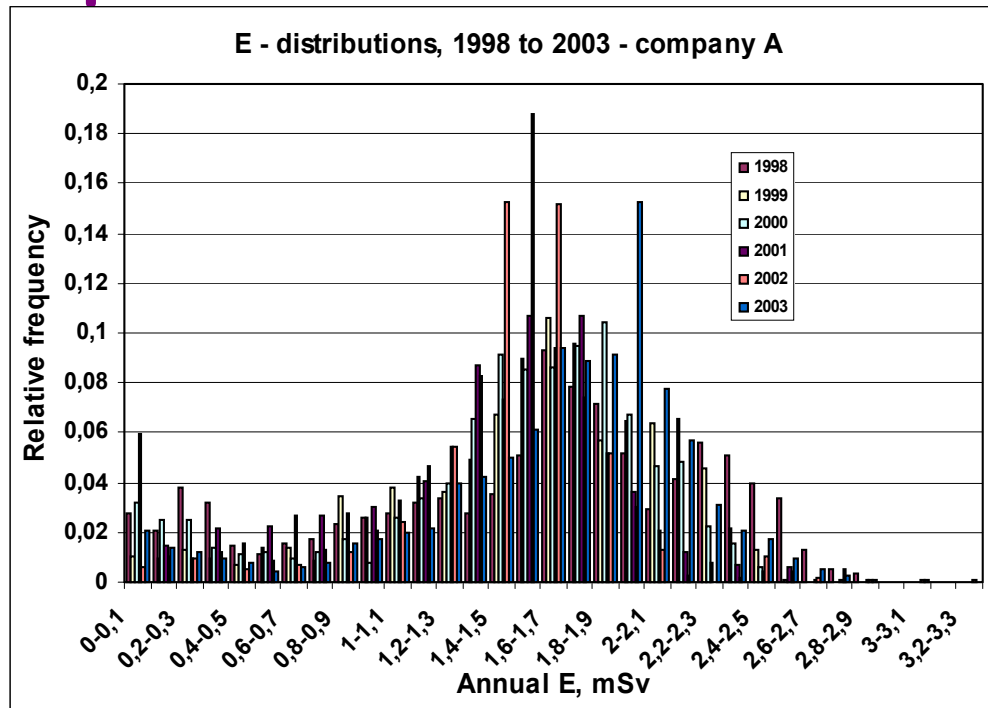
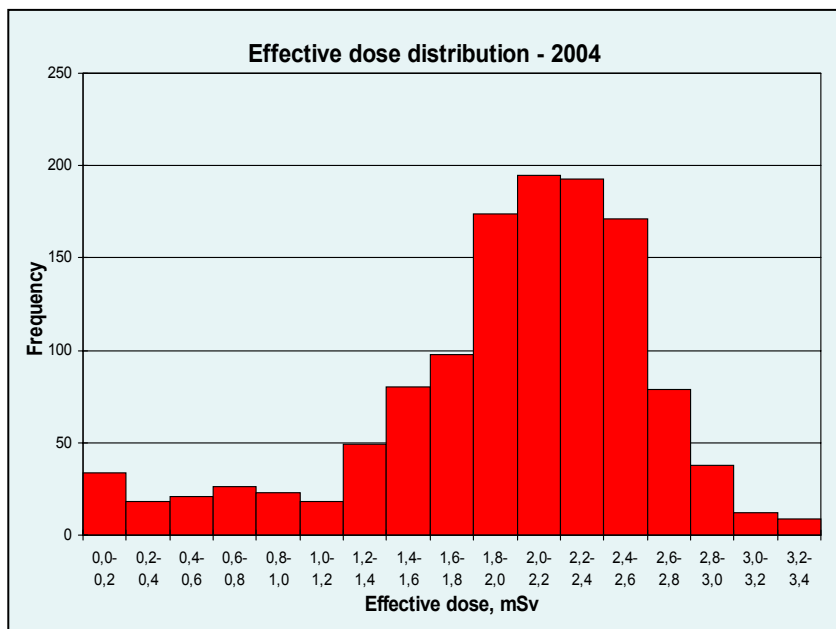
$\text{CaSO}_4 : \text{Dy} : (1.5 - 1.8) \text{ mSv}$

ROUTINE INDIVIDUAL DOSIMETRY PROCEDURE

- 1. Air company representative prepares the data sets characterising the aircraft and persons traffic for a period**
- 2. Sets are used to create an input set for the CARI code**
- 3. The calculation of individual effective dosis for each flight performed, simplified flight profile:**
 - only actual parameter: flight time**
 - calculated for the model routes parameters, i.e.**
 - flight altitude ; and the time of the rise and the falling down**
- 4. The effective dose of each aircrew member are calculated and transmitted to an air company**

REMARKS: All data since 1998 now recalculated by CARI 6
Generally – more than 90 % above 1 mSv

Individuel dosimetry of aircraft crew - czech air companies



External Exposure to Natural Radiation

Results of some Experimental Studies

2 - Onboard spacecraft exposure

Dosimetry and microdosimetry methods

- **Thermoluminescent detectors (TLD's)**
 - $\text{Al}_2\text{O}_3\text{:C}$; LiF:Mg,Cu,P ; $\text{CaSO}_4\text{:Dy}$; $> 1 \mu\text{Gy}$
 - AIP glass; LiF:Mg,Ti ; $> 10 \mu\text{Gy}$

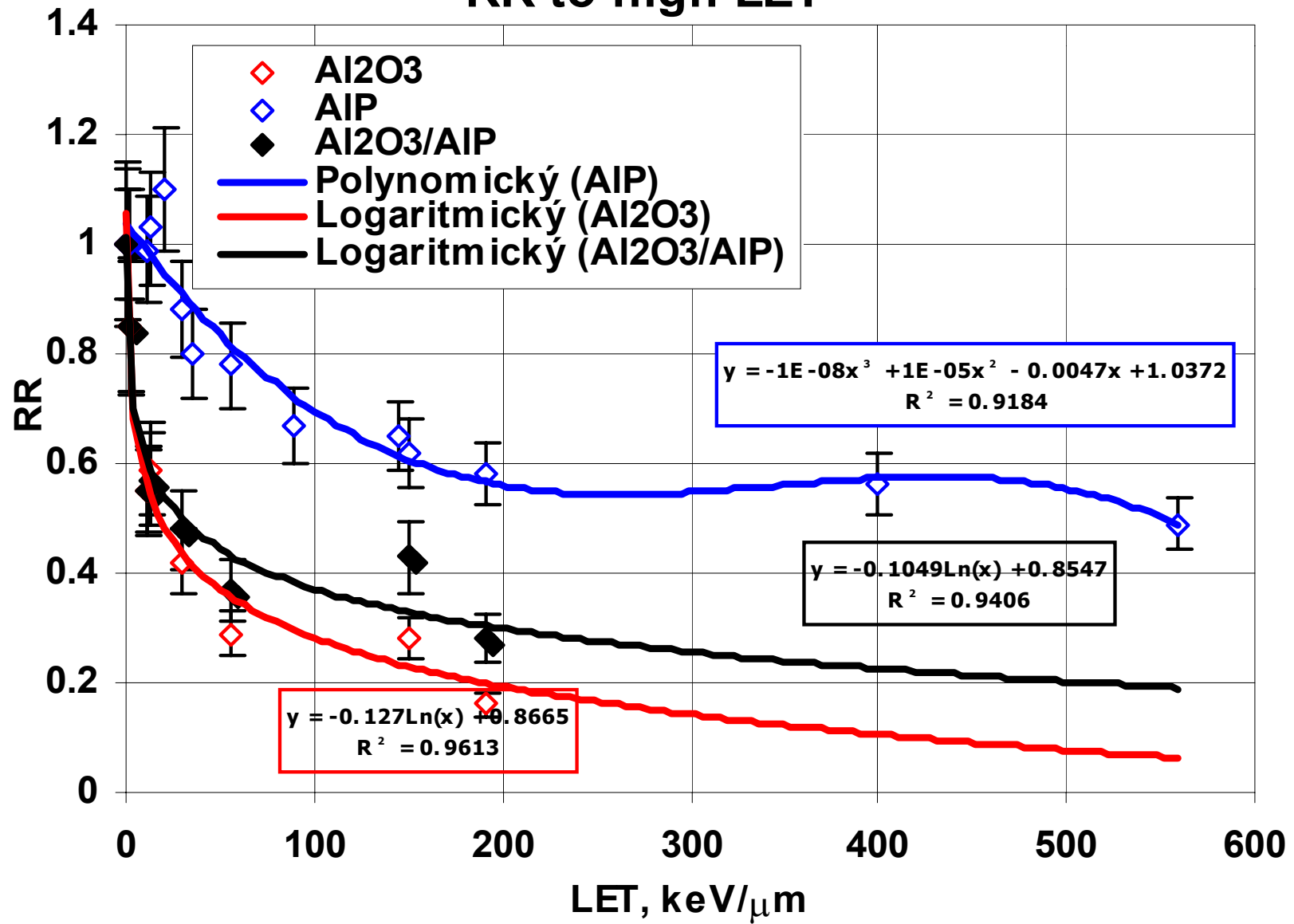
Different $\text{RR} = f(\text{LET})$
- **Spectrometer of the LET** - H and D distributions in $\text{LET} \in (\approx 10; 700) \text{ keV}/\mu\text{m}$ in tissue; 1 to 100 mSv

Experiments and analysis 2004-2005

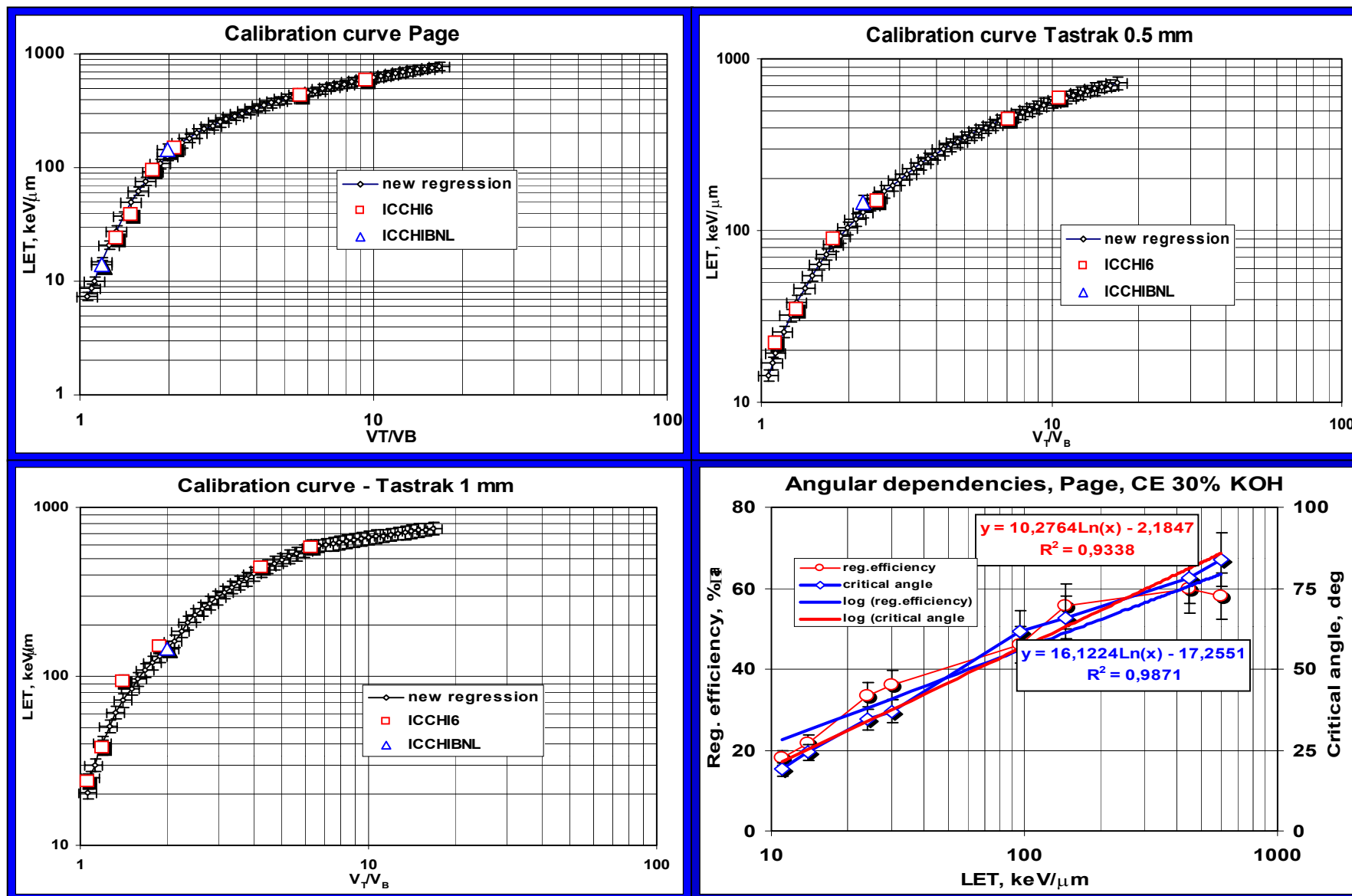
On-Earth's calibrations

- **ICCHIBAN 6** (C, Ar, Kr; 24 - 600 keV/ μm)
- **ICCHIBAN-NSRL** (H, O, Fe; 0.2 – 150 keV/ μm)
- **ICCHIBAN 8** (He, O, Ar, Fe; 2.2 – 320 keV/ μm)
- **Dubna Nuclotron** (C, Mg, Fe; 8 – 200 keV/ μm)

RR to high LET



Calibration of TED for LET spectrometry; angular dependence of the response



Some more recent onboard SS measurements

Mission	Period	Shield g.cm⁻²	Altitude km	TLD μGy/day	D-TED μGy/day	H-LET μGy/day
MIR 28	6/04/– 16/06/00	~ 15	330–350	140 ± 10	13 ± 1	85 ± 5
ISS-1	30/11/01– 3/11/02	~ 20	380–420	212 ± 15	22 ± 2	202 ± 12
MESSAGE	17/10/– 28/10/03	~ 15	380–420	166 ± 8	16 ± 2	223 ± 22
ISS-2	31/01/– 31/10/04	~ 15	340-360	150 ± 8	15 ± 1	98 ± 5

Interpretation of measured data

Made supposing that:

- TLD's data characterize mostly of radiation with LET below few keV per μm (primary protons);
- the ratio of high LET secondary particle dose and primary proton collision dose is equal to (0.025 ± 0.003) ;
- the contribution of primary long range cosmic heavier charged particles represents about 22% of total LET spectrometer established dose, resp. 34% of LET spectrometer established dose equivalent value.

Full dosimetric characteristics onboard of space stations

Mission	D, $\mu\text{Gy/day}$		Neutrons in % of D	H, $\mu\text{Sv/day}$		Neutrons in % of H
	>10 keV/ μm	total		>10 keV/ μm	total	
MIR 28	17 ¹⁾	157	6.0	129	269	22
ISS-1	28	240	7.0	306	518	27
MESSAGE	21	186	6.4	338	504	34
ISS-2	19	191	5.6	148	322	22

¹⁾ 1 S.D. estimated to 15-20%

Comparison

MDU-Liulin spectra

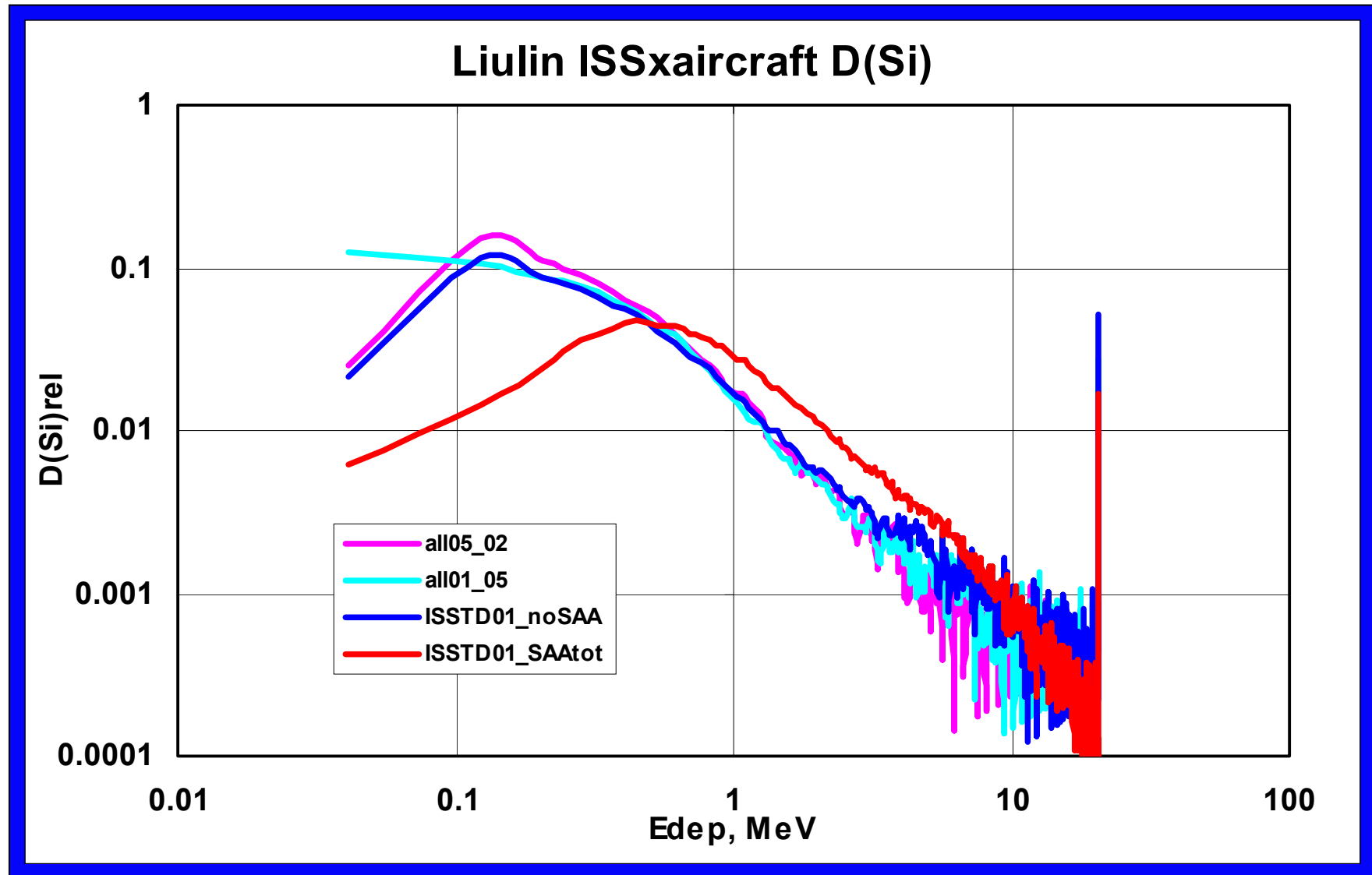
onboard aircraft

(2001, 2005 –in total ~ 200 μ Sv);

and onboard ISS

(2001 – no SAA ~ 70 μ Gy; in SAA ~ 360 μ Gy)

Comparison of Si-energy deposition spectra



Comparison

*MDU-Liulin spectra onboard aircraft (2001, 2005 ~ 200 μ Sv);
and onboard ISS (2001 - no SAA ~68.6 μ Gy; in SAA ~ 360 μ Gy)*

When supposed that:

- neutron spectra onboard ISS and aircraft are similar;
- relative excess above 1 MeV onboard ISS is due to HECP of GCR with QF ~ 5; and
- the correction factors for non-neutron and neutron-like component are the same onboard ISS as onboard aircraft;

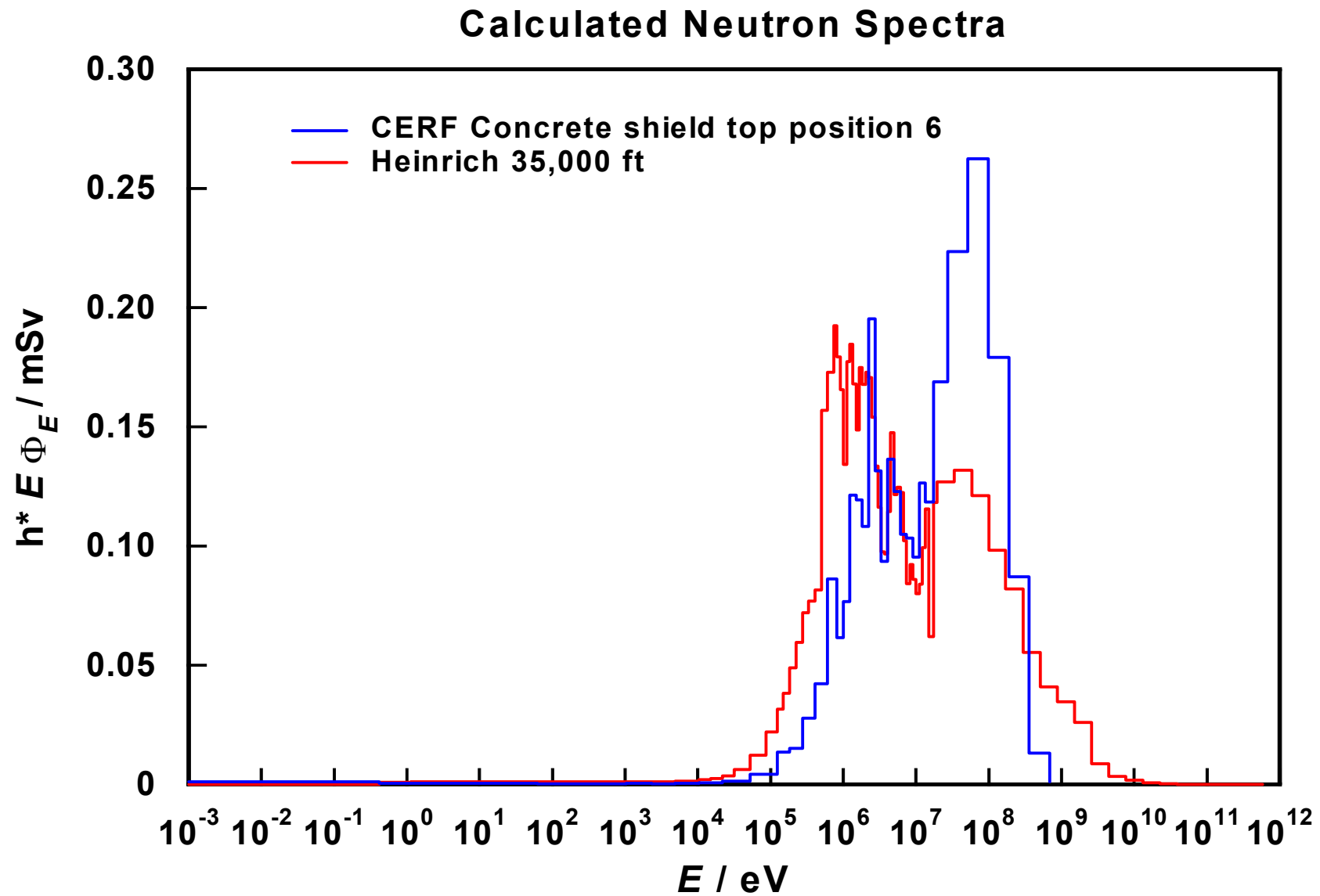
Than outside SAA the daily values of “H*(10)” are:

- non-neutron component: 72.3 μ Sv;
- neutron-like component: 119.3 μ Sv;
- HECP component: 64.6 μ Sv; and
- Total: 256 μ Sv
- Average QF ~ 2.8

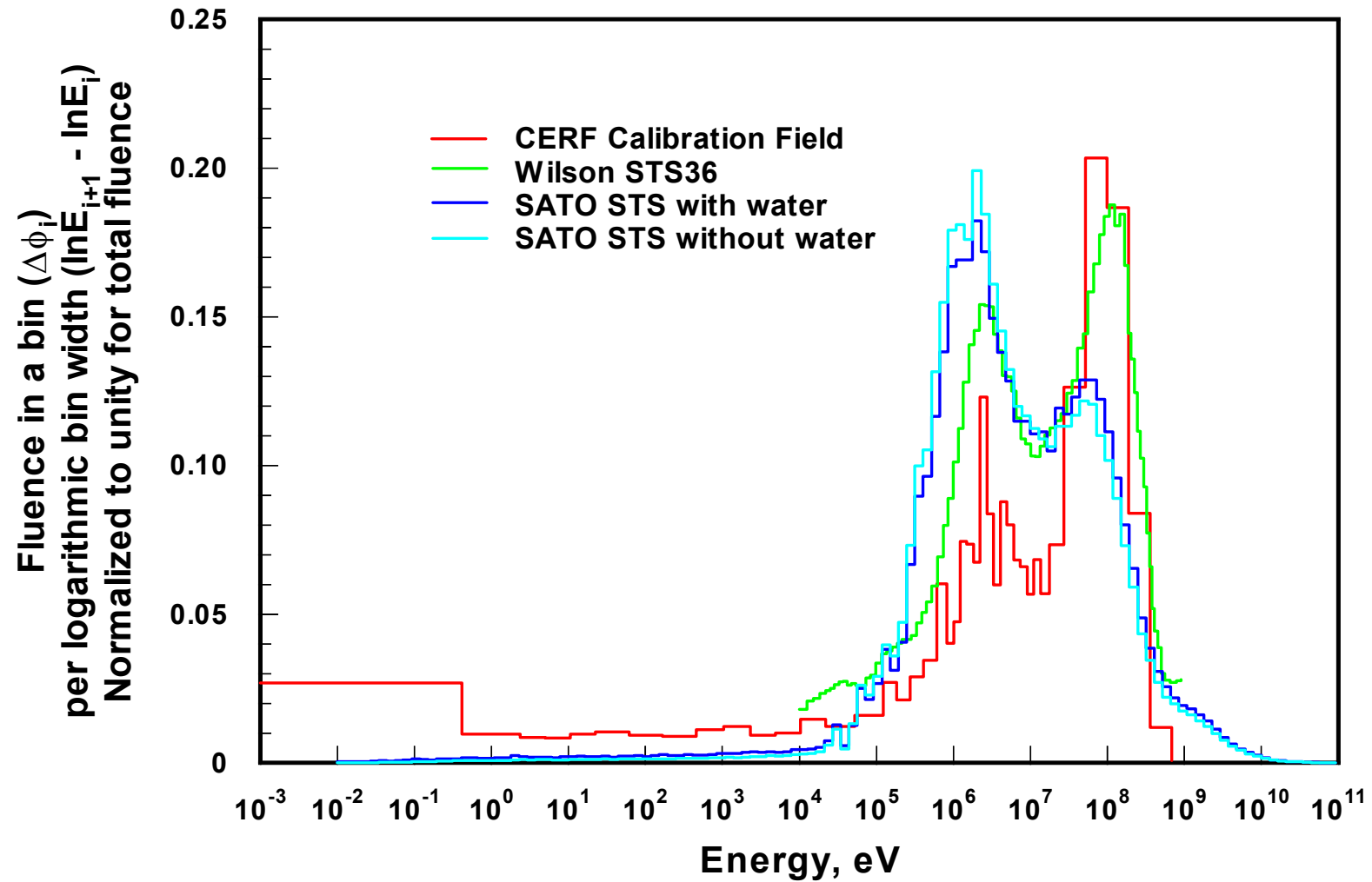
When in SAA mostly protons with QF ~ 1.5 than

Total ~ 610 μ Sv; QF ~ 2.0

Comparison of CERF and at 35kfeet calculated spectra



Further comparison of neutron spectra



Experiments going on; future - also with TLDs INRNE

- Long term exposure (>3 years) – ISS;
- Further on-Earth calibrations
- Onboard ISS dosimetry and LET spectrometry:
 - Interior of a phantom (russian Matrjoshka);
 - In-area distribution of dosimetry characteristics
 - Further comparison MDU on ISS/aircraft
- DOBIES project (ESA); since 2006

Basic references

***Transport Methods and Interactions for Space Radiation.* NASA Reference Publication 1257; eds.: Wilson, J.W. et al., NASA, December 1991**

***1990 Recommendations of the International Commission on Radiological Protection.* ICRP Publications 60, Annals of ICRP 21, (1-3), 1991**

***Exposure of Aircrew to Cosmic Radiation.* EURADOS Rep.1996-01, eds.: I.R.McAulay et al., ISBN 92-827-7994-7, Luxembourg 1996**

W. Friedberg et al.: *Computer Program CARL.* US Department of Commerce, NTS, Springfield, VA, 1994

***Operational Radiation Safety Program for Astronauts in low Earth Orbit: A Basic Framework.* NCRP Rep. No. 142, Bethesda, 2002**

***Dosimetry of Aircrew Exposure to Radiation During Solar Maximum.* Final Report, EC Contract No. FIGM-CT-2000-00068, Coordinator: Denis O'Sullivan, DIAS, Dublin 2004**

***Cosmic Radiation Exposure of Aircraft crew: Compilation of Measured and Calculated data.* Editors: L.Lindborg, D. Bartlett, P. Beck, I. R. McAulay, K. Schnuer, H. Schraube and F. Spurný, EC-DG TREN, ISBN 92-894-8448-9, Luxembourg 2004**

G. Reitz: *Radiation. Limiting Human spaceflight?* Keynote Lecture, 15th Humans in Space Symposium, Graz, May 22-26, 2005

D. Bartlett: private communication, 2005

Thank you for your attention !